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ASTROPHYSICS

HARVARD & SMITHSONIAN

# FLARE STATISTICS AND HIGH-RESOLUTION SPECTROSCOPY OF A VOLUME COMPLETE SAMPLE OF MID-TO-LATE M DWARFS WITHIN 15 PARSECS

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Lucile &  
**Packard**  
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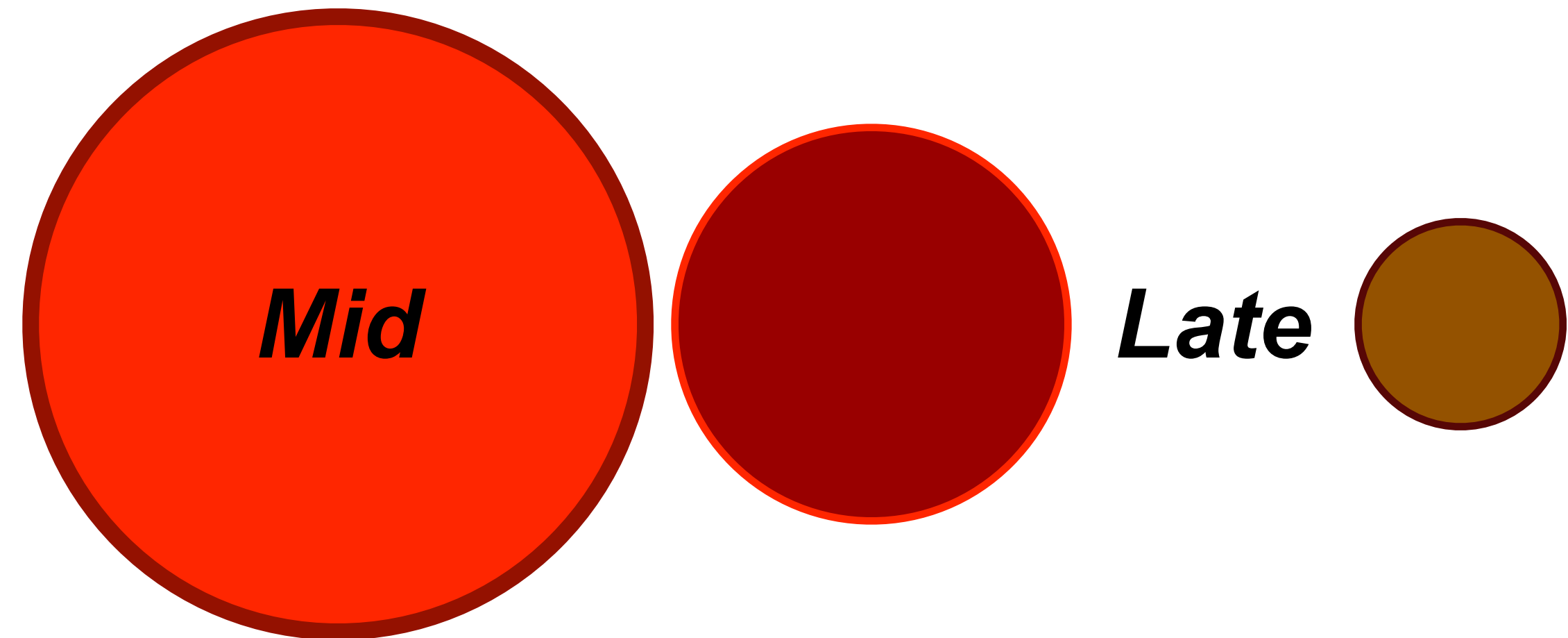
 **JOHN  
TEMPLETON**  
FOUNDATION



## MID-TO-LATE M DWARFS

$$0.1 M_{\text{Sun}} < M < 0.3 M_{\text{Sun}}$$

- Are fully convective
- Lack a tachocline
- Have magnetic activity
- Magnetic field generation may be different in low-mass stars



### **M4V**

$$M = 0.25 M_{\text{sun}}$$

$$R = 0.25 R_{\text{sun}}$$

$$T = 3100 \text{ K}$$

### **M6V**

$$M = 0.15 M_{\text{sun}}$$

$$R = 0.15 R_{\text{sun}}$$

$$T = 2800 \text{ K}$$

### **M8V**

$$M = 0.1 M_{\text{sun}}$$

$$R = 0.1 R_{\text{sun}}$$

$$T = 2600 \text{ K}$$

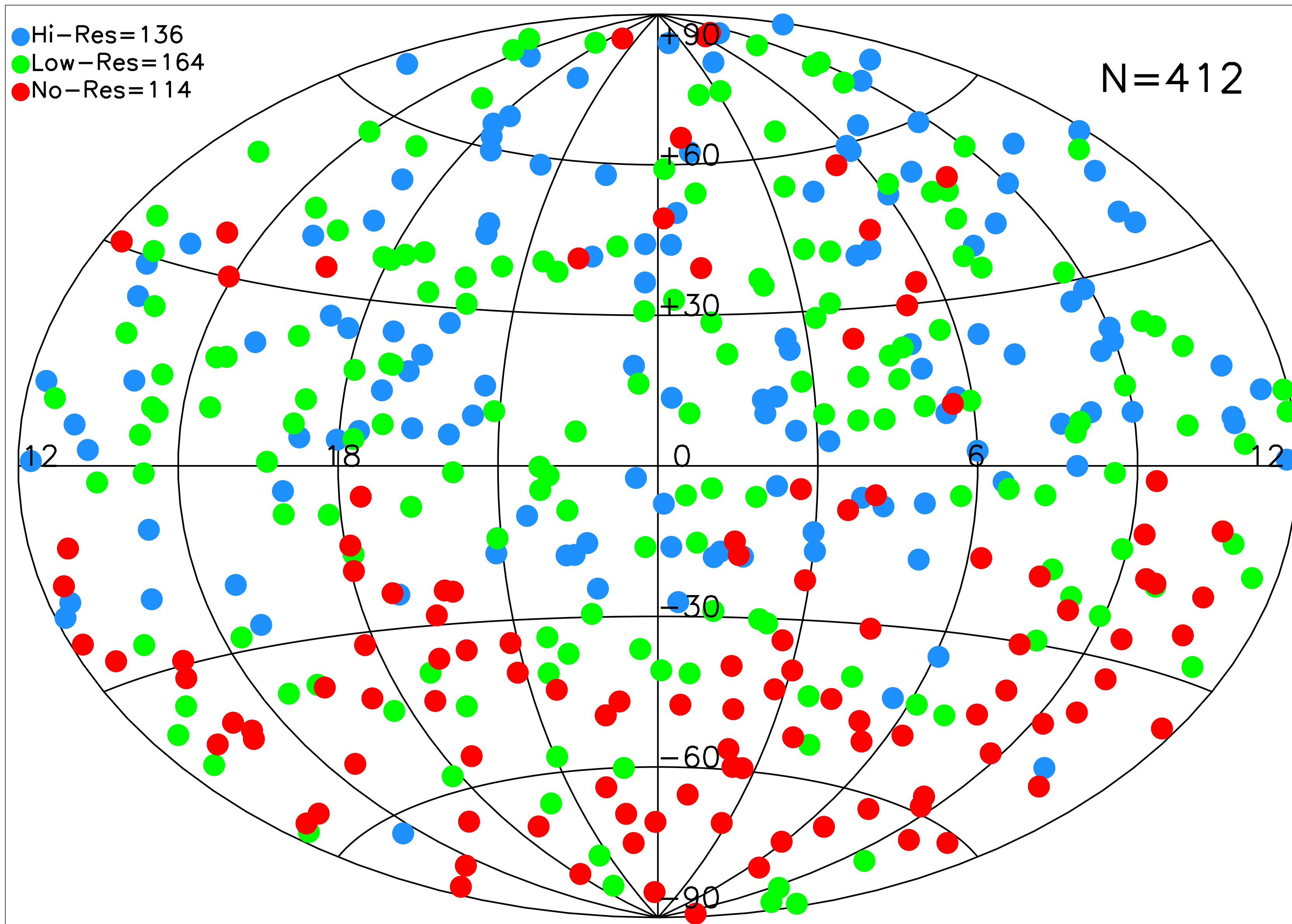
**HOW DOES THE RATE OF LARGER, ENERGETIC FLARES  
RELATE TO STELLAR ROTATION AND ACTIVITY?**

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# VOLUME COMPLETE SAMPLE OF MID-TO-LATE M DWARFS WITHIN 15 PARSECS

- We are gathering multi-epoch high-resolution spectra of each star.
- In the south, we use the CHIRON Spectrograph located on the SMARTS 1.5 meter telescope at Cerro Tololo Inter-American Observatory, Chile.
- In the north, we use the TRES Spectrograph located on the 60 inch telescope at Fred Lawrence Whipple Observatory on Mt. Hopkins, Arizona.

- We are
- In the 1.5 m
- In the telescope Arizona

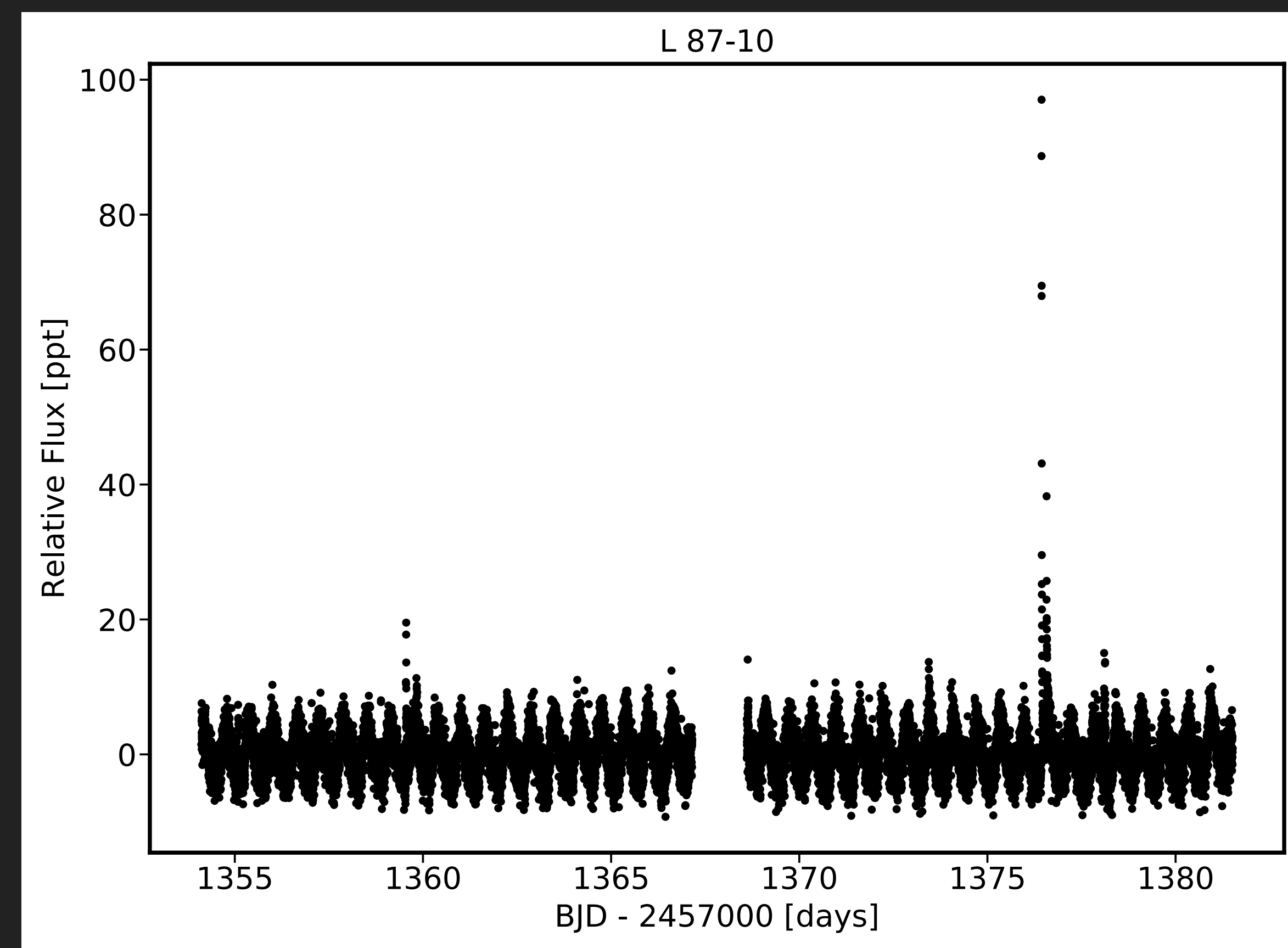
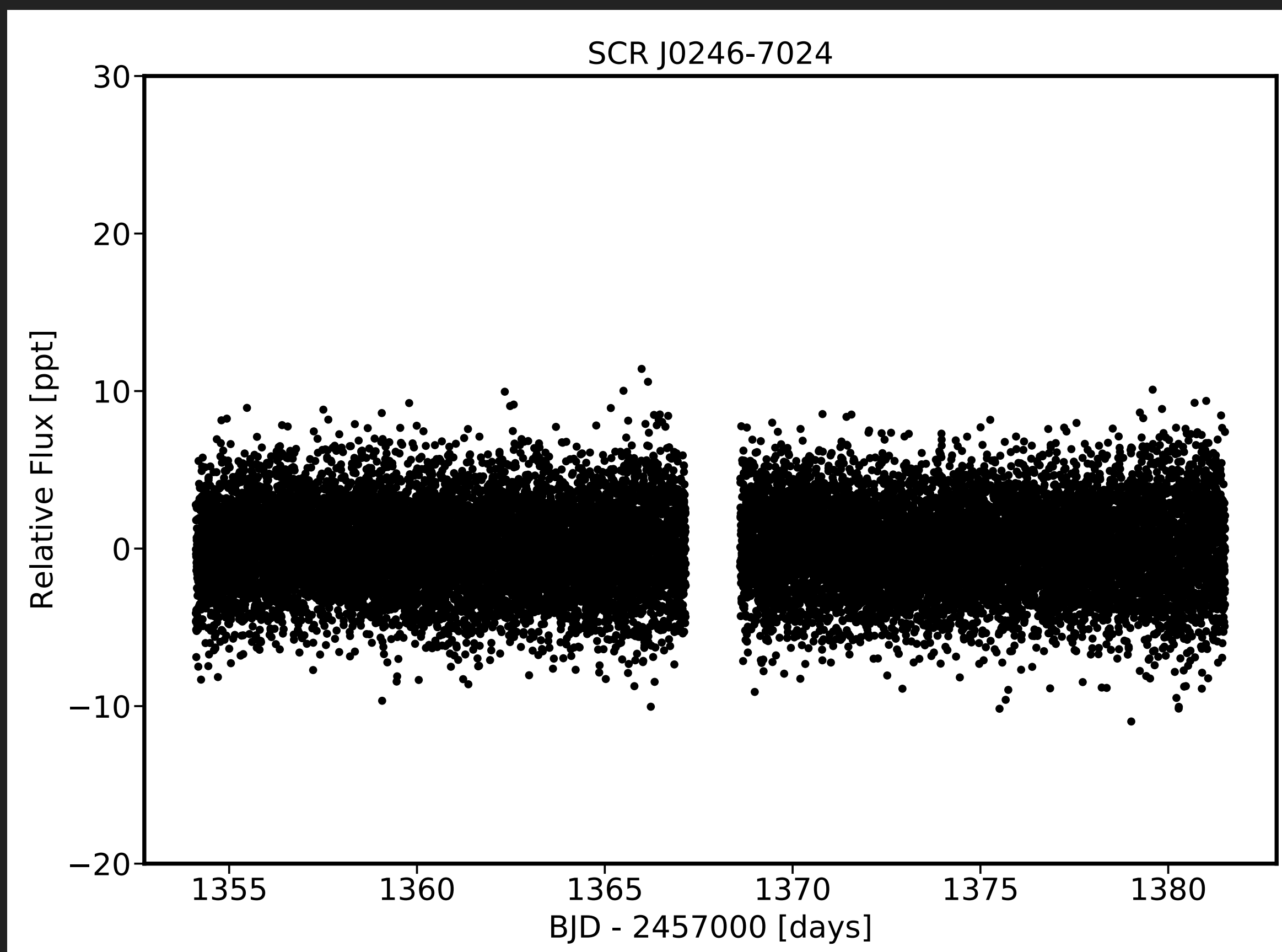


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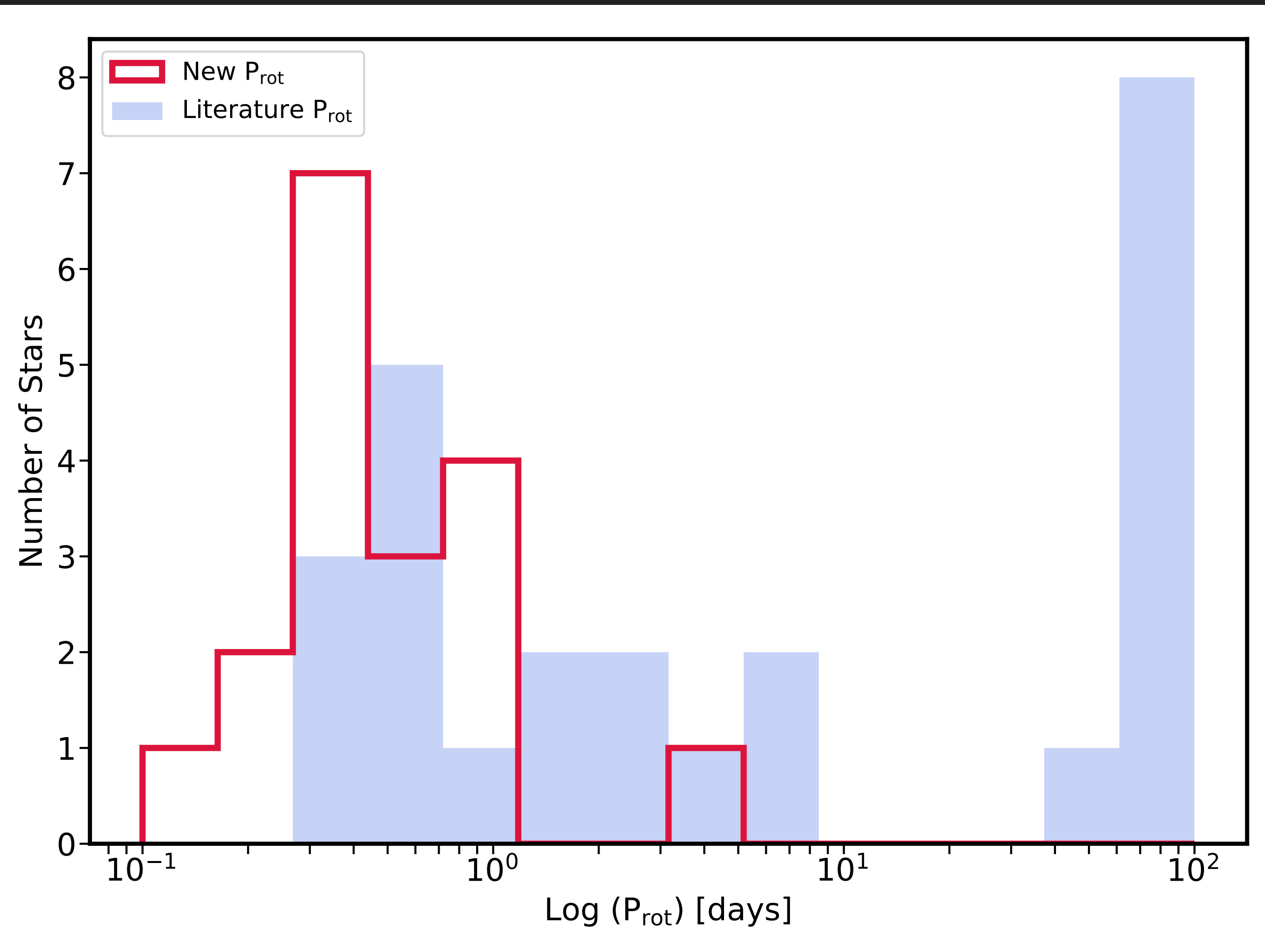
# TESS DATA OF NEARBY MID-TO-LATE M DWARFS!

2 minute cadence TESS data off mid-to-late M dwarfs within 15

parsecs! TESS GI Proposals: G011231, G022076 | PI Jennifer Winters



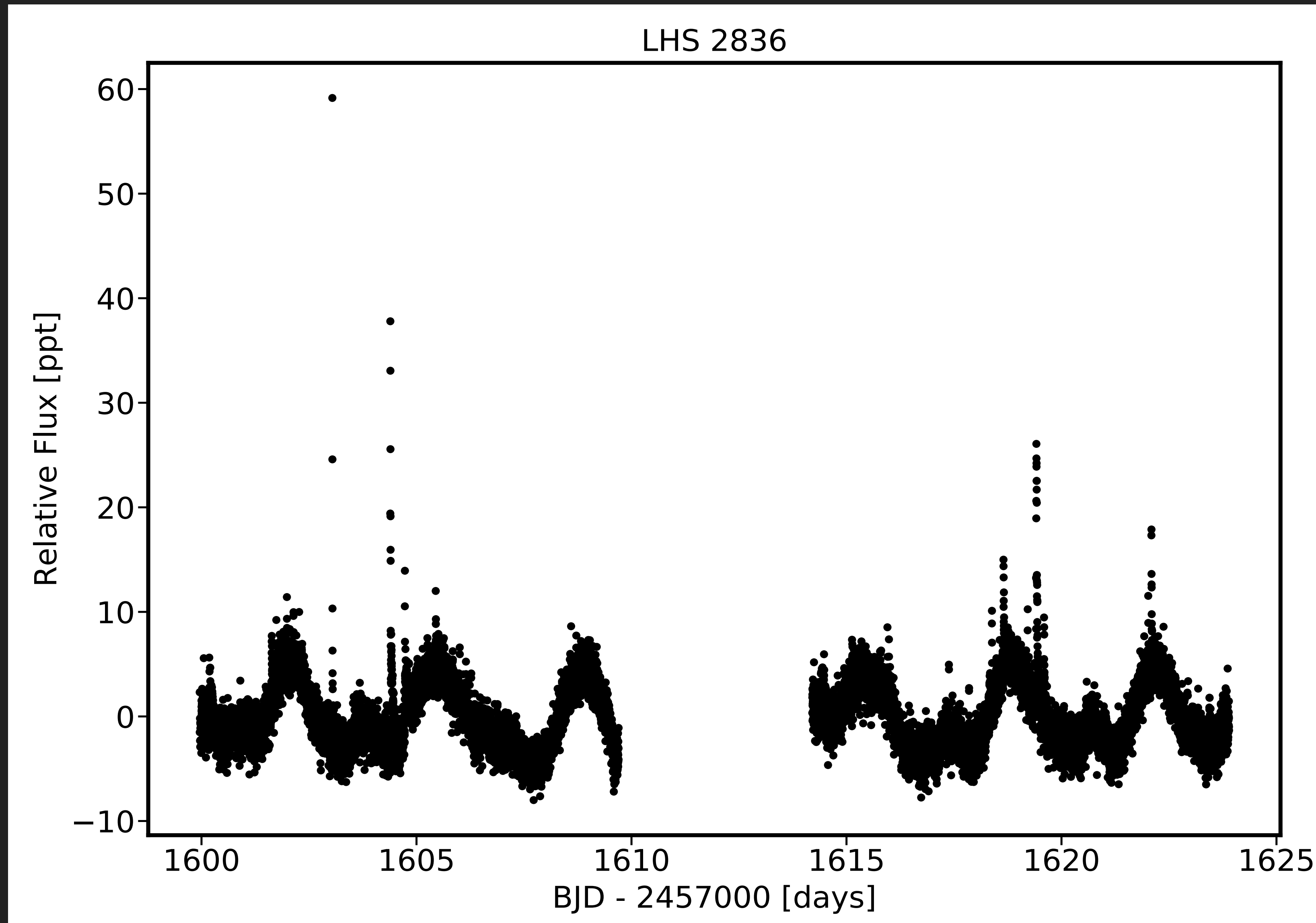
# NEARBY MID-TO-LATE M DWARFS PROVIDE A RICH HUNTING GROUND FOR FLARES WITH TESS!



- 154 stars have been observed in sectors 1-12.
- 116 are single stars.
- 36 stars have a previously published rotation period.
- We measured 17 new rotation periods with TESS.

# FINDING FLARES IN TESS DATA

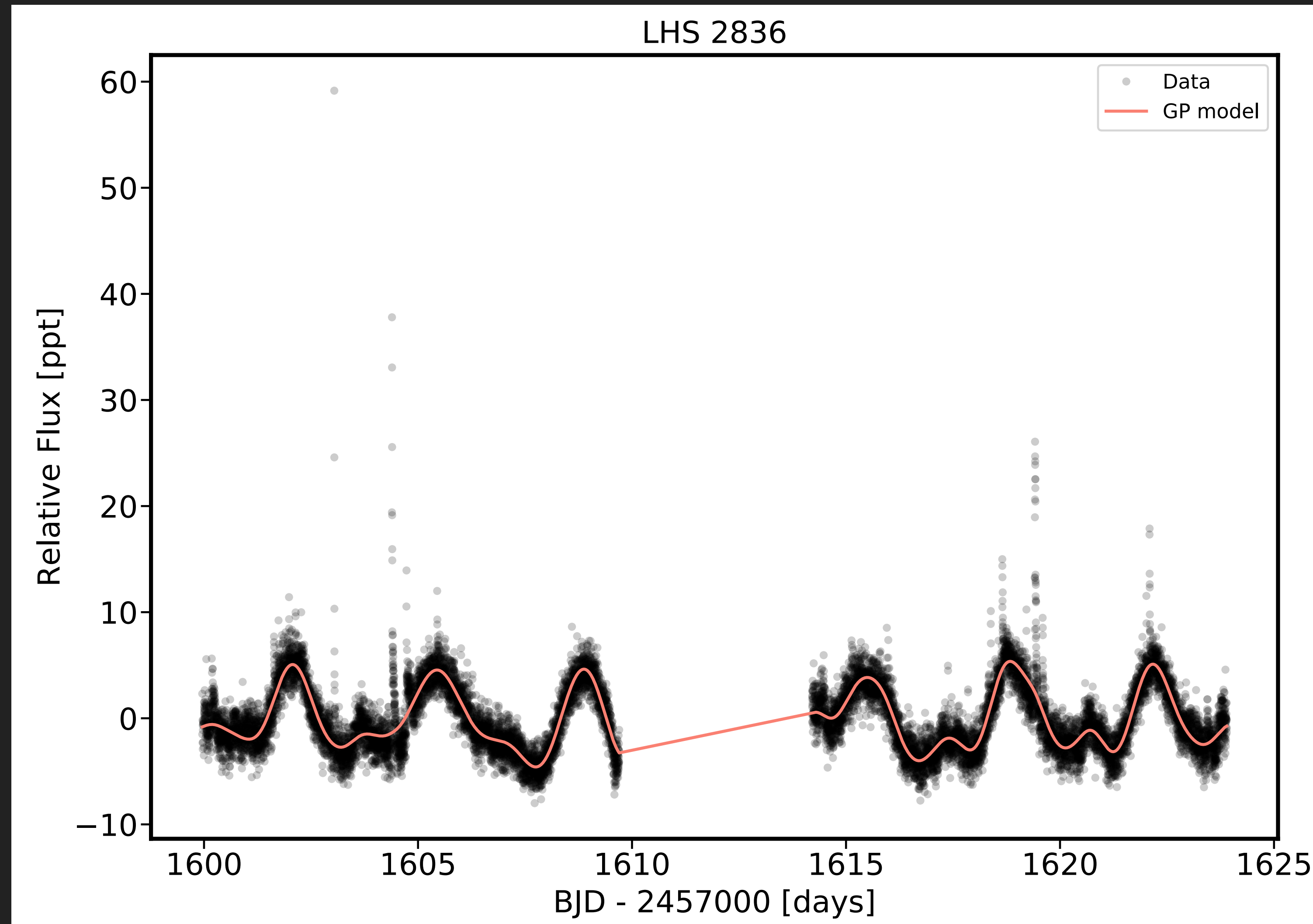
- ▶ We use PDC TESS light curve





# FINDING FLARES IN TESS DATA

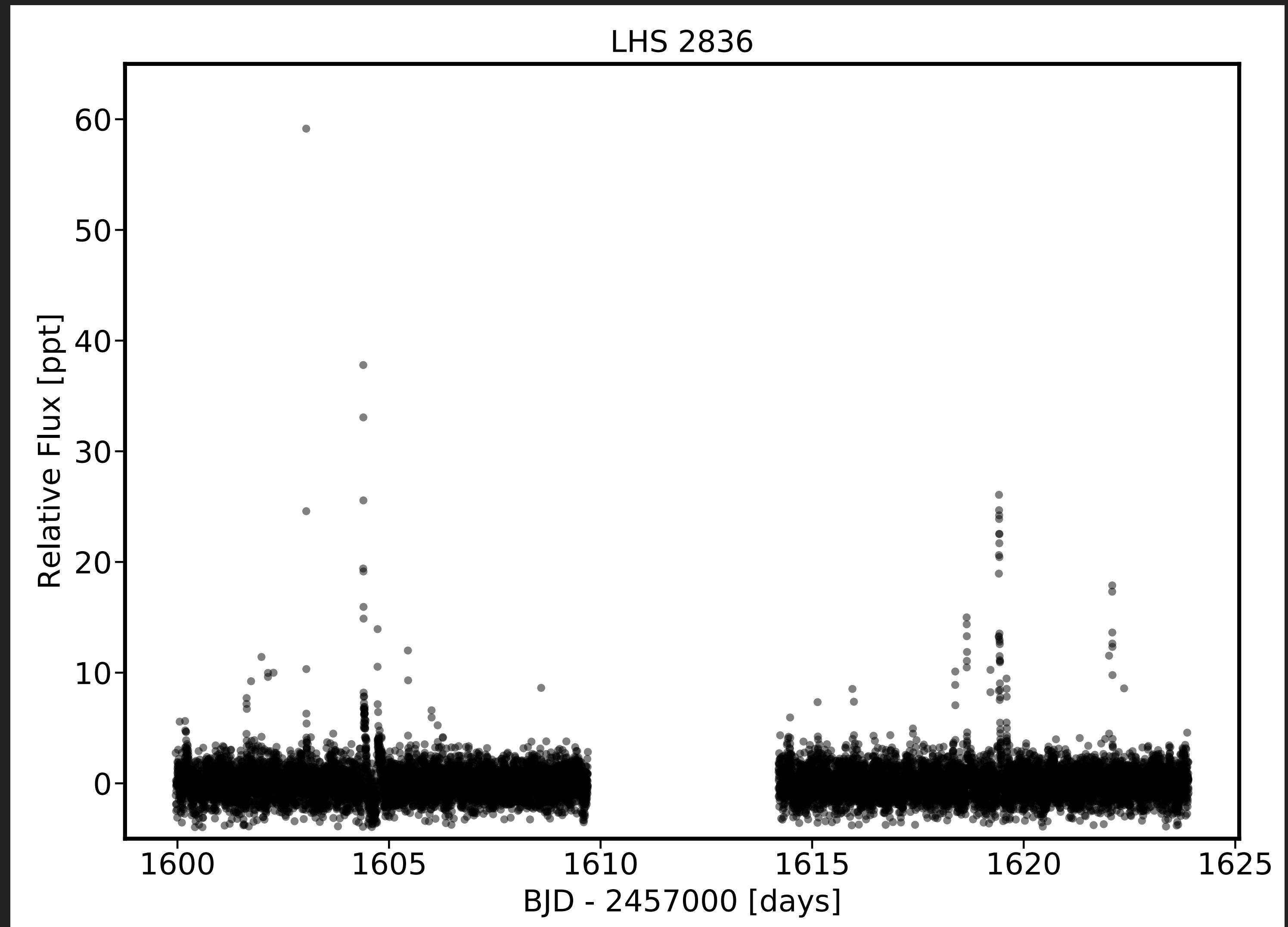
- ▶ We use a Gaussian Process model to de-trend the data.



# FINDING FLARES IN TESS DATA

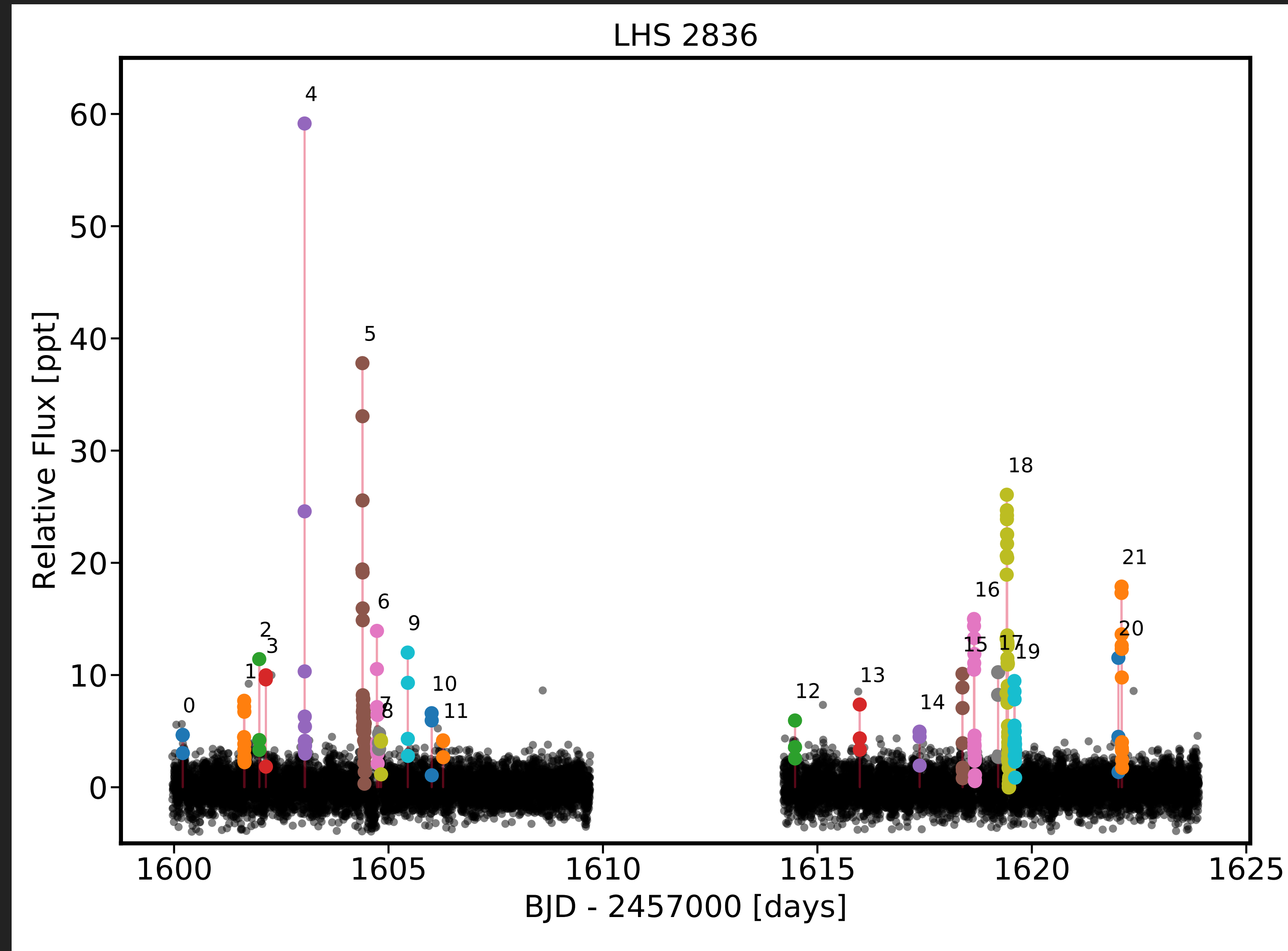
- ▶ We search for flares using the de-trended light curves.

- ▶ The flares are detected using the criteria from Chang et al. (2015)



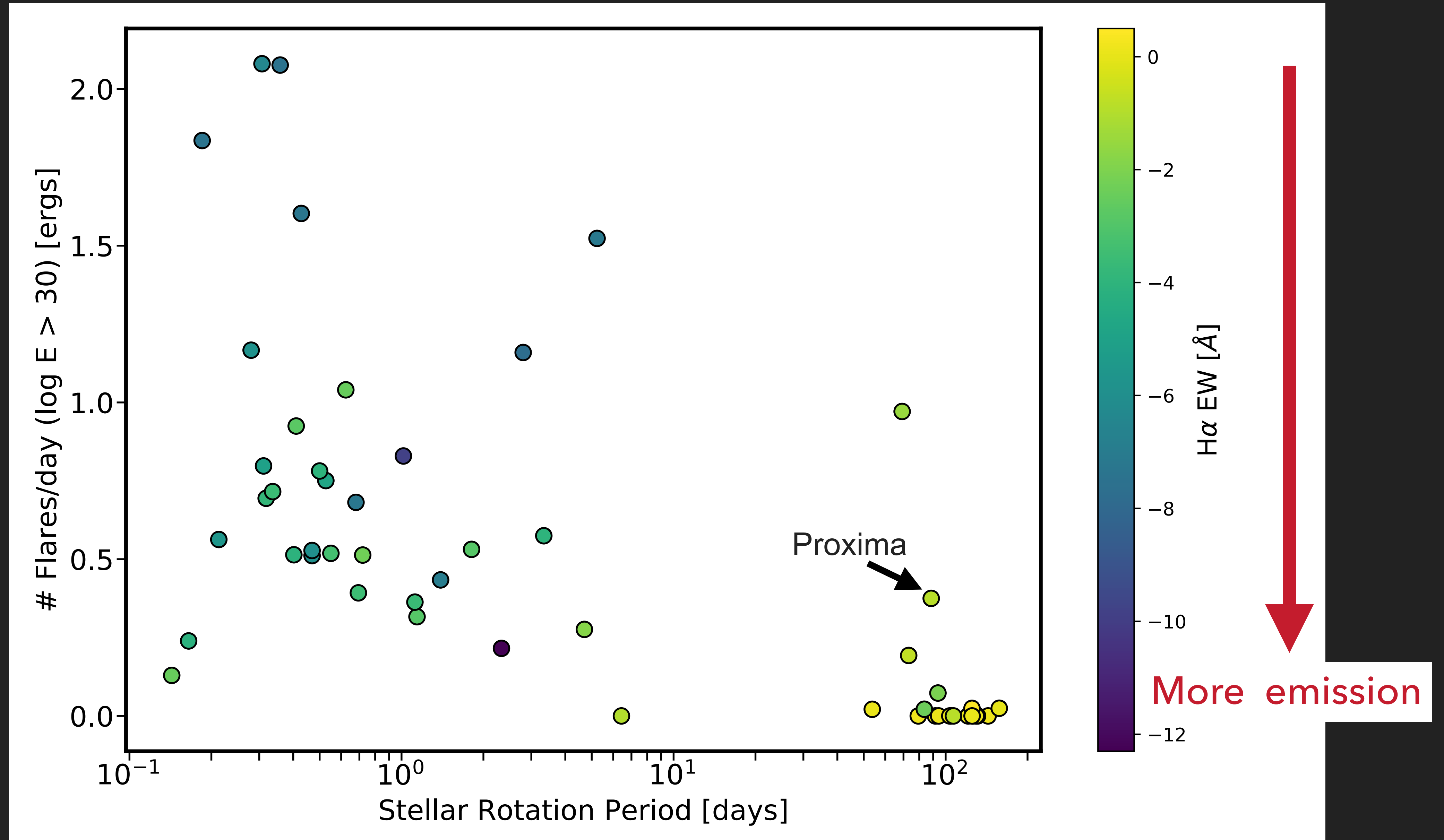
# FINDING FLARES IN TESS DATA

- ▶ We search for flares using the de-trended light curves.





# FLARE RATES OF MID-TO-LATE M DWARFS



# SUMMARY

- ▶ We are measuring flare rates of all  $(0.1 - 0.3) M_{\text{Sun}}$  mid-to-late M dwarfs within 15 parsecs observed by TESS.
  - ▶ In sectors 1-12, 36 stars have published rotation periods.
  - ▶ We measure 17 new rotation periods.
- ▶ Stars with  $P_{\text{rot}} < 85$  days show H $\alpha$  in emission and a high rate of energetic flares. This rate does not correlate with rotation period.
- ▶ Stars with  $P_{\text{rot}} > 100$  days show no H $\alpha$  in emission and little to no energetic flares.
- ▶ More to come with continuing analysis of sectors 13-26!